

XMASS experiment

WIN05

8th June 2005

A. Takeda for the XMASS collaboration
Kamioka Observatory, ICRR,
University of Tokyo

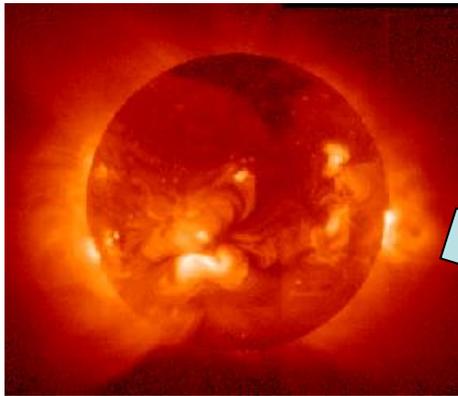
1. Introduction
2. R&D status using prototype detector
3. Summary

1. Introduction

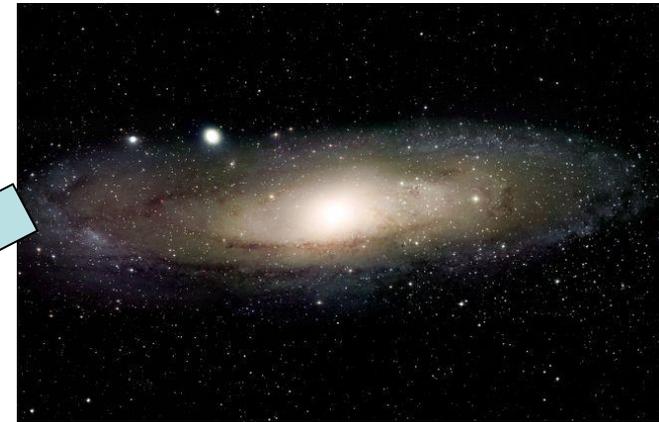
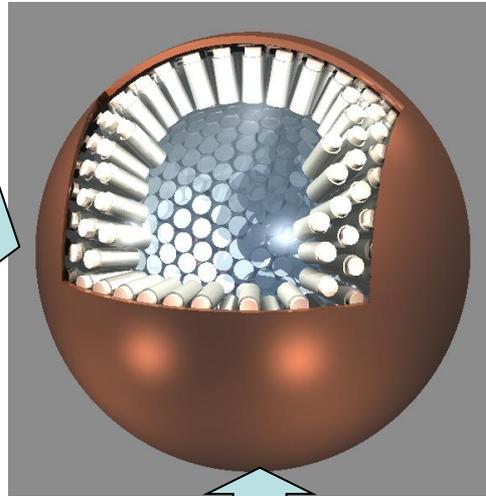
➤ What's XMASS

Multi purpose low-background experiment with liq. Xe

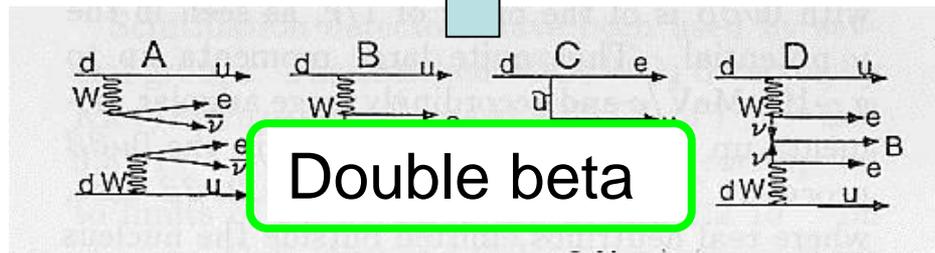
- **X**enon **MASS**ive detector for solar neutrino (**pp**/⁷**Be**)
- **X**enon neutrino **MASS** detector (**ββ decay**)
- **X**enon detector for Weakly Interacting **MASS**ive Particles (**DM search**)



Solar neutrino



Dark matter

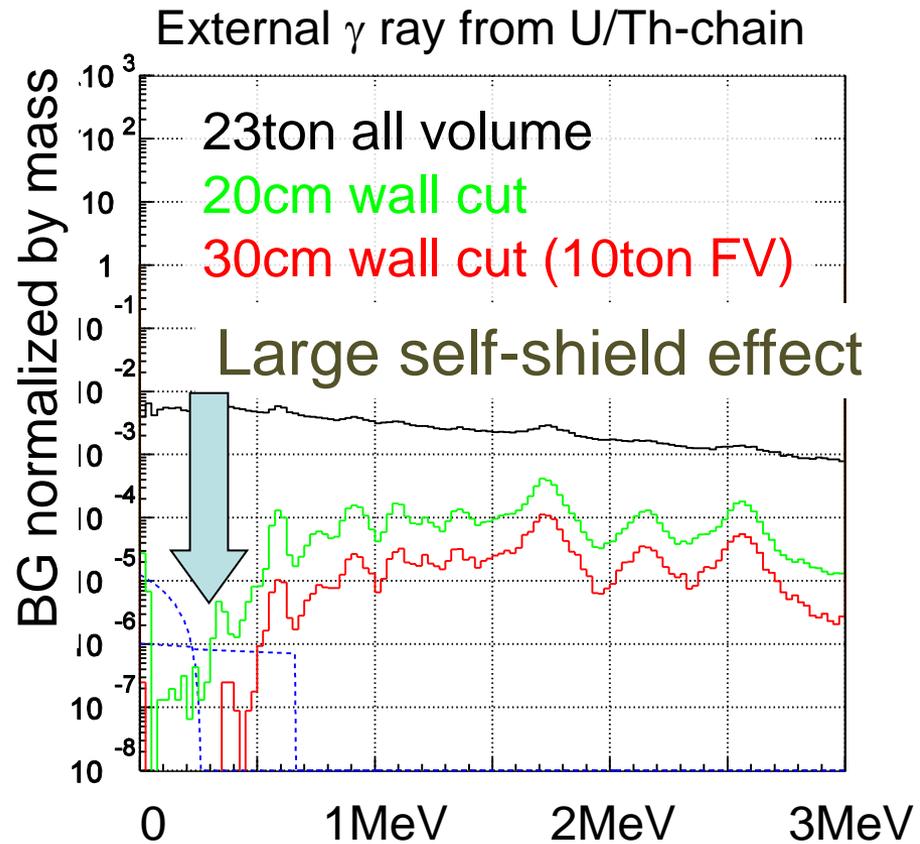
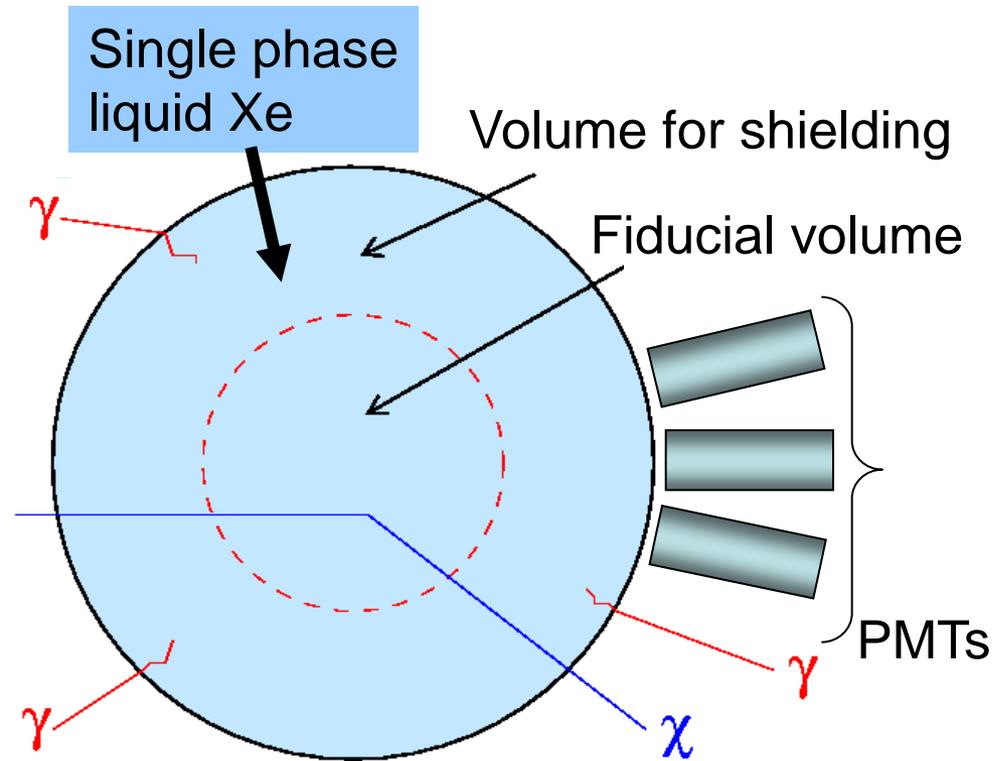


➤ Why liquid xenon

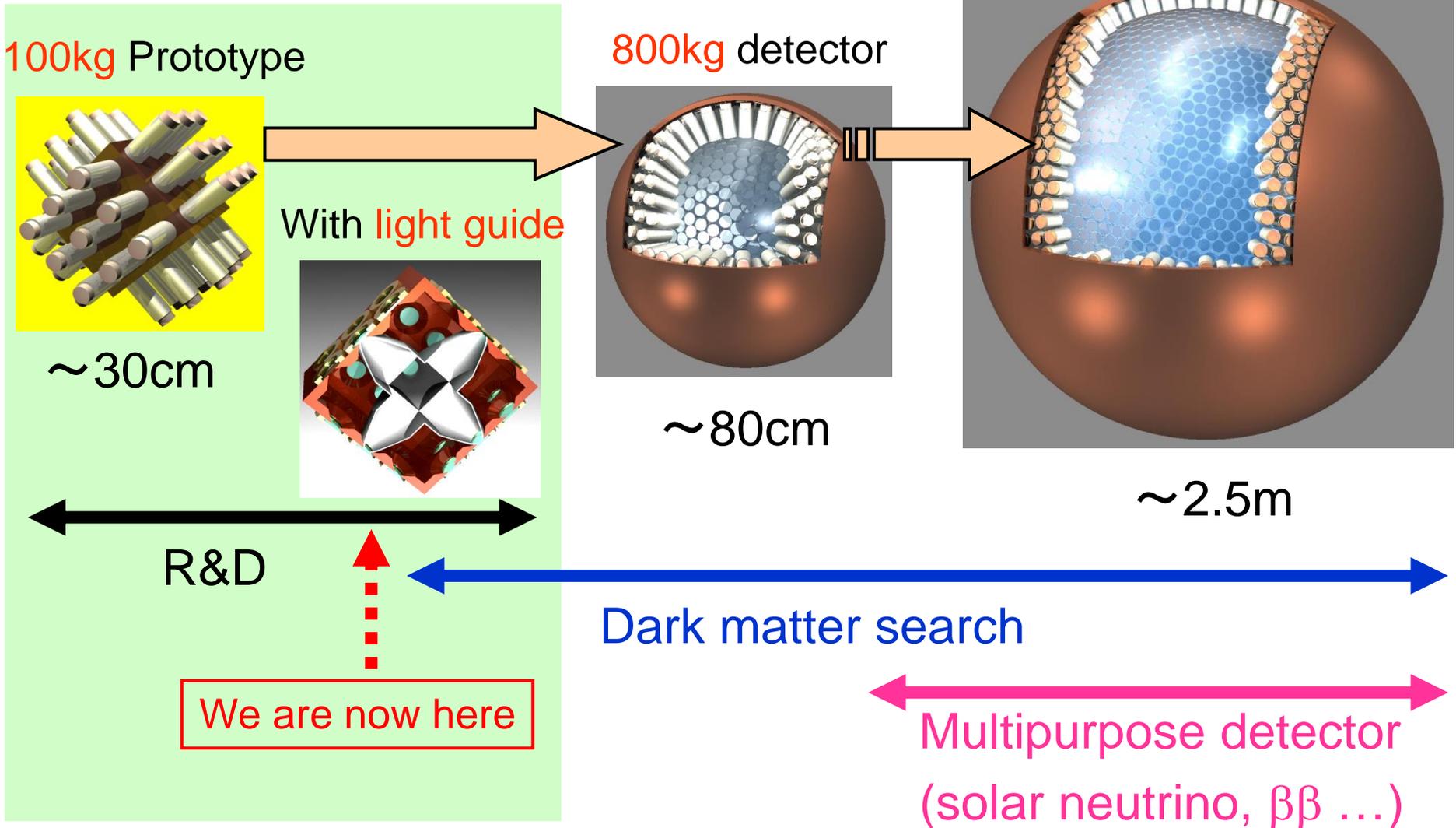
- **Large Z (=54)**
Self-shielding effect
- **Large photon yield (~42 photons/keV ~ NaI(Tl))**
Low threshold
- **High density (~3 g/cm³)**
Compact detector (10 ton: sphere with diameter of ~2m)
- **Purification (distillation)**

- No long life radioactive isotope
- Scintillation wavelength (175 nm, detected directly by PMT)
- Relative high temperature (~165 K)

➤ Key idea:
self-shielding effect for low energy events

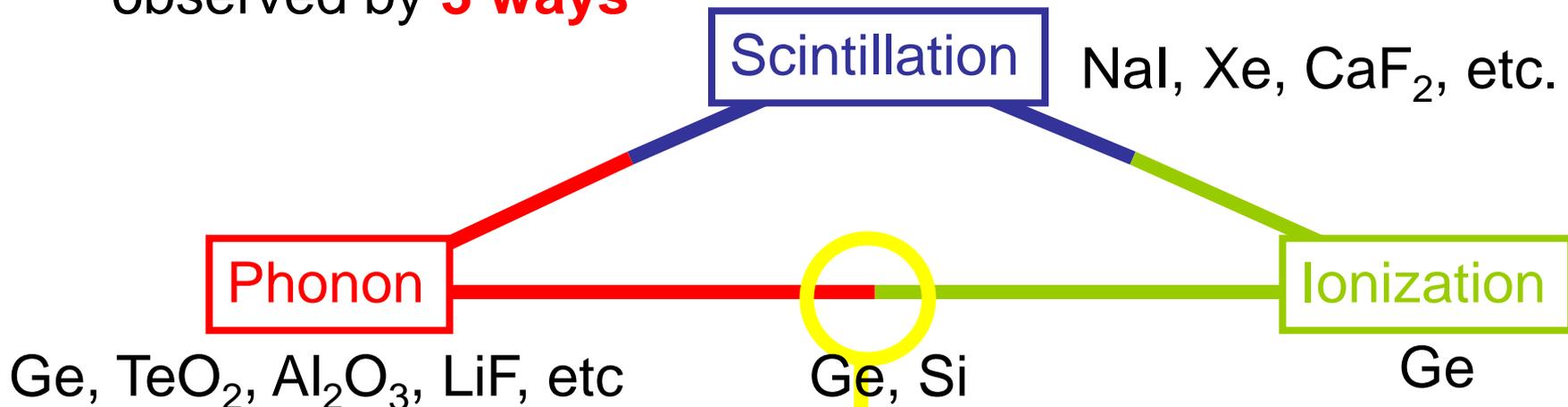


➤ Strategy of the scale-up



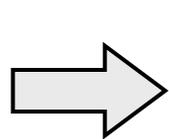
Trend of Dark matter (WIMPs) direct searches

- Recoiled nuclei are mainly observed by **3 ways**



- Taking **two type of signals simultaneously** is recent trend

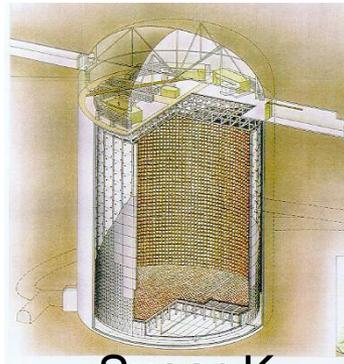
CDMS, EDELWEISS: phonon + ionization



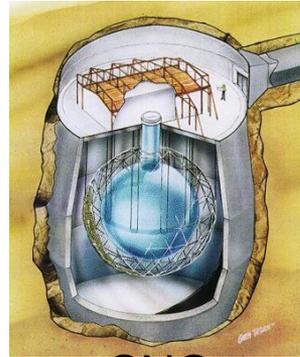
- ◆ γ ray reduction owing to powerful particle ID
- ◆ However, seems to be difficult to realize a large and uniform detector due to complicated technique

Strategy chosen by XMASS

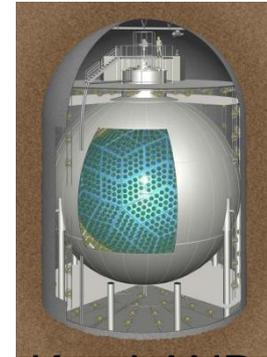
- Make **large mass** and **uniform** detector (with liq. Xe)
 - ➔ Same style as successful experiments of Super-K, SNO, KamLAND, etc.
- Reduce γ ray BG by fiducial volume cut (**self shielding**)



Super-K



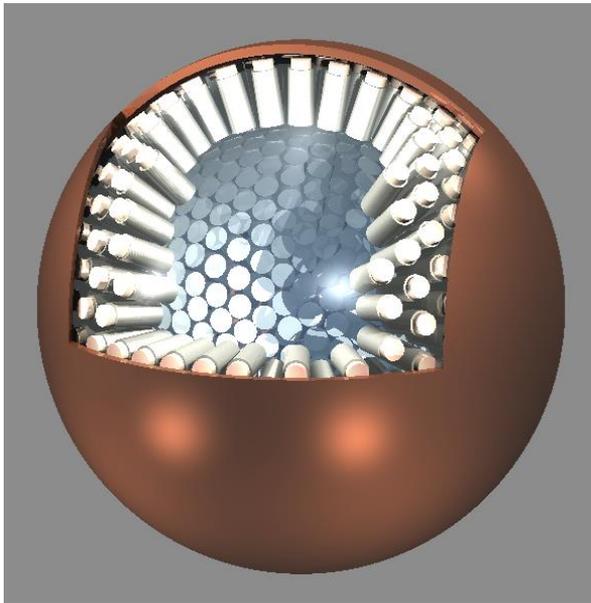
SNO



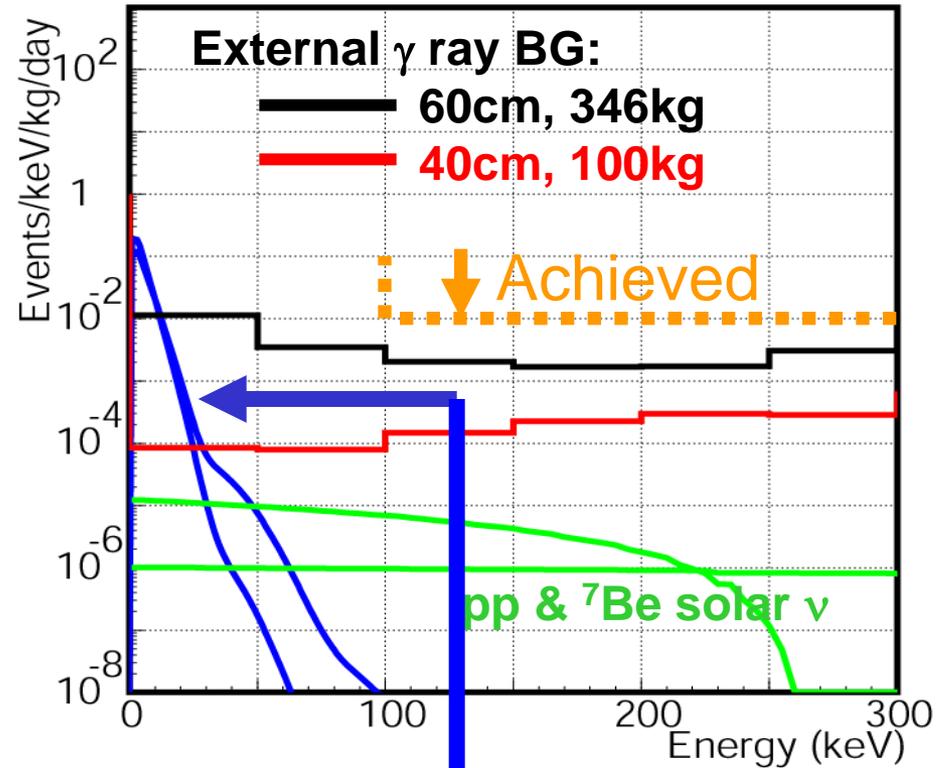
KamLAND

➤ 800 kg detector

Main purpose: **Dark Matter search**



~80cm diameter



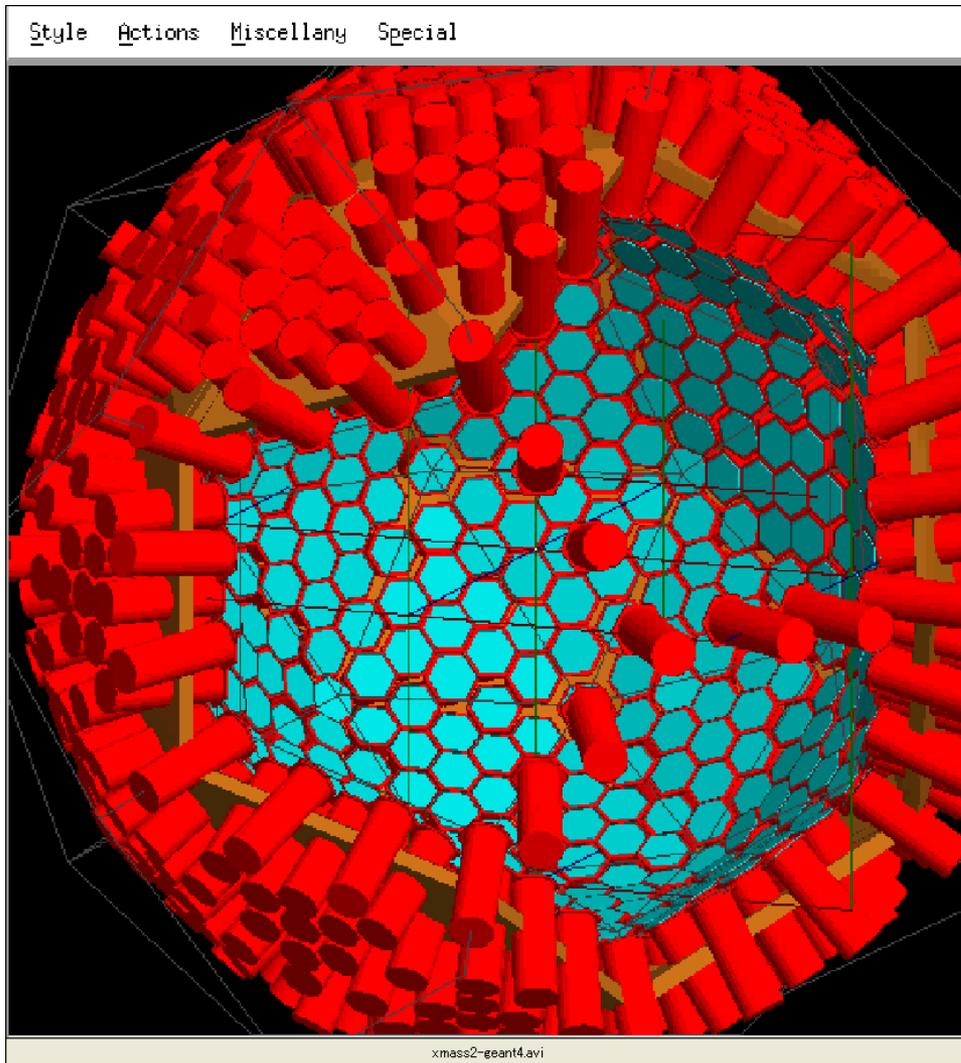
- **~800-2" PMTs**
immersed into liq. Xe
- **70%** photo-coverage



~5 keVee
threshold

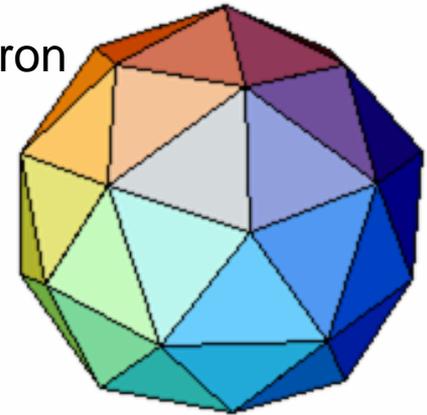
Expected dark matter signal
(assuming 10^{-42} cm^2 , Q.F.=0.2
50GeV / 100GeV,)

Geometry design



- A tentative design (not final one)

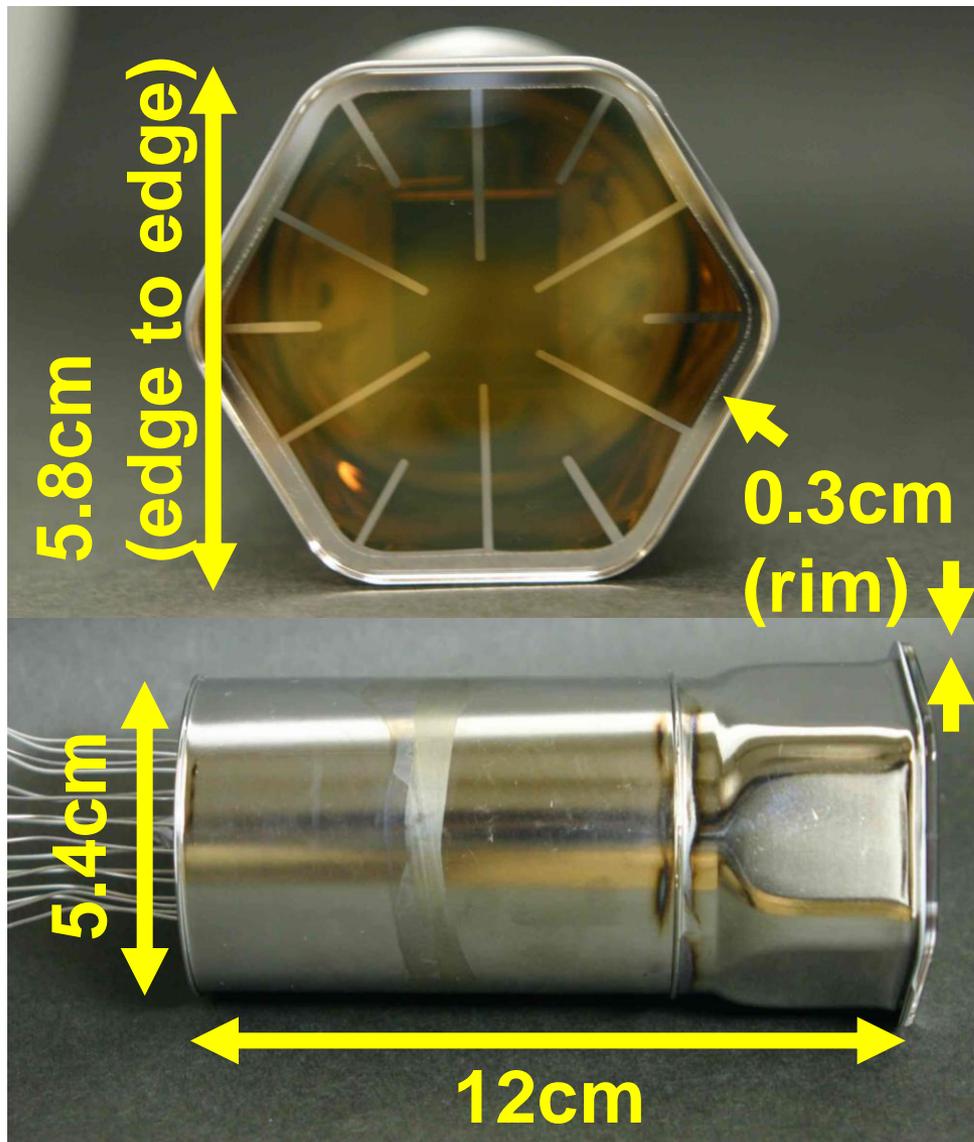
12 pentagons /
pentakisdodecahedron



- Total **840** hex **PMTs** immersed into liq. Xe
- **70%** photo-coverage
- Radius to inner face **~43cm**

▲ This geometry has been coded in a Geant 4 based simulator

Hamamatsu R8778MOD(hex)



- Hexagonal quartz window
- Effective area: $\phi 50\text{mm}$ (min)
- QE $< \sim 25\%$ (target)
- Aiming for 1/10 lower background than R8778

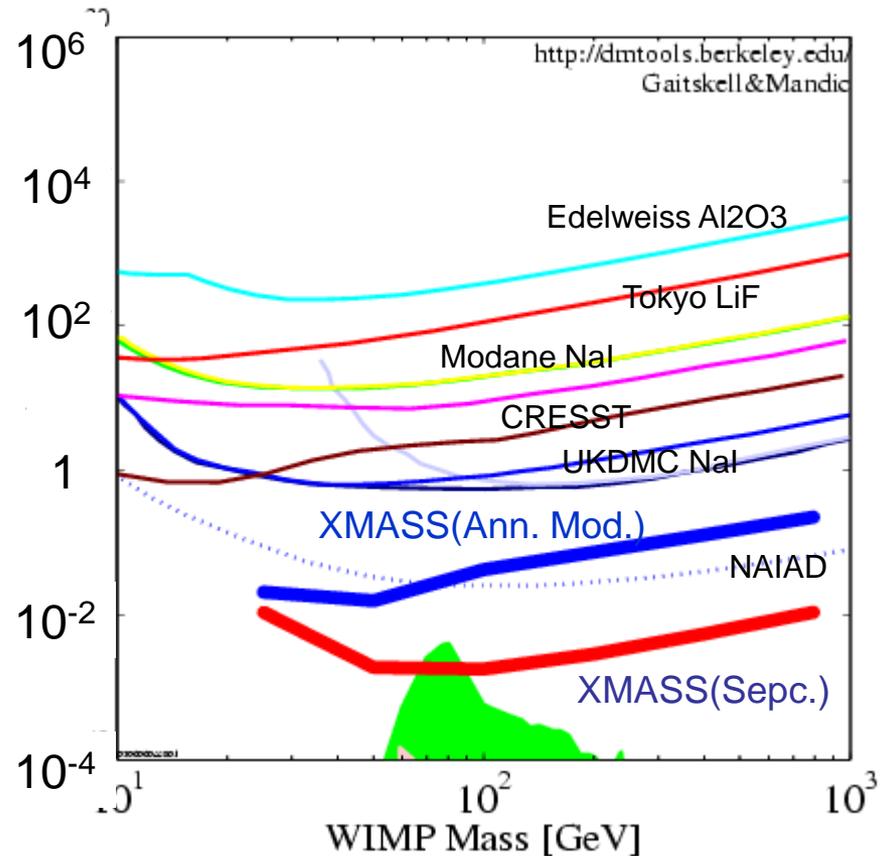
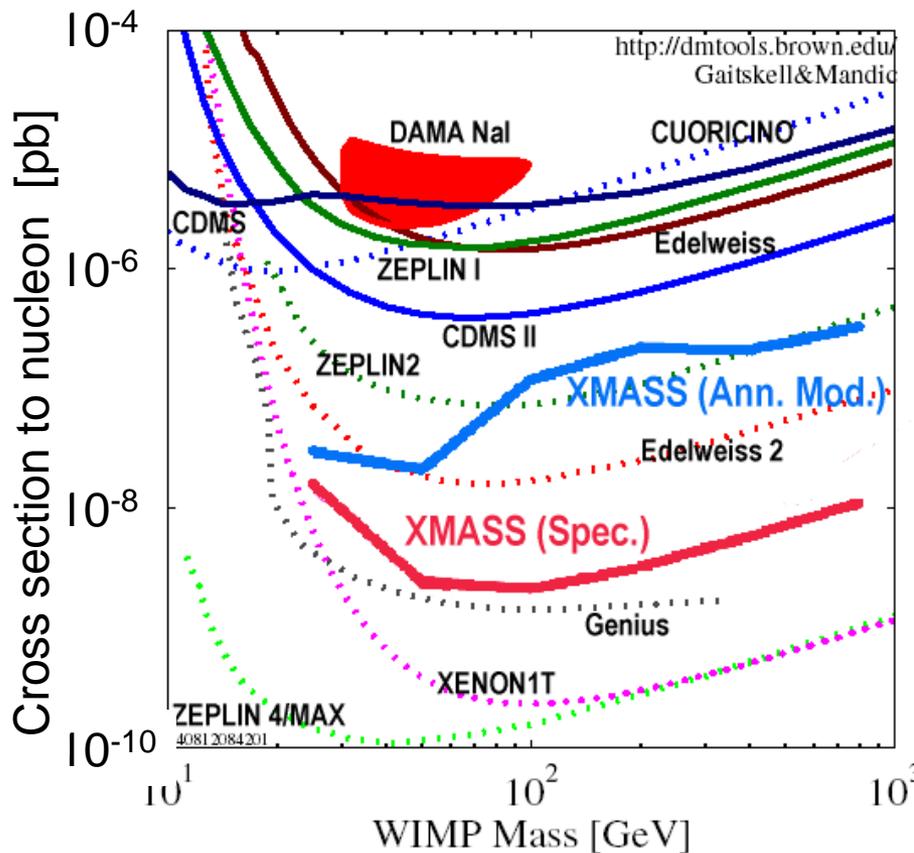
c.f. R8778

U	$1.8 \pm 0.2 \times 10^{-2}$ Bq
Th	$6.9 \pm 1.3 \times 10^{-3}$ Bq
^{40}K	$1.4 \pm 0.2 \times 10^{-1}$ Bq

- **Prototype has been manufactured already**
- Now, being tested

Expected sensitivities

XMASS FV 0.5 ton year
 $E_{th} = 5 \text{ keVee} \sim 25 \text{ p.e.}$, 3σ discovery
 w/o any pulse shape info.



- Large improvements will be expected
 $SI \sim 10^{-45} \text{ cm}^2 = 10^{-9} \text{ pb}$
 $SD \sim 10^{-39} \text{ cm}^2 = 10^{-3} \text{ pb}$

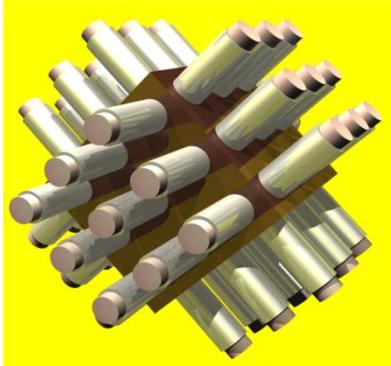
Plots except for XMASS:
<http://dmtools.berkeley.edu>
 Gaitskell & Mandic

2. R&D status using prototype detector

➤ Main purpose

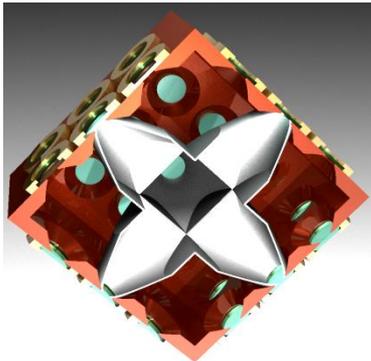
- Confirmation of estimated 800 kg detector performance
 - ◆ **Vertex and energy reconstruction** by fitter
 - ◆ Miss fitting due to dead angle of the cubic detector (“**wall effect**”, will be explained later) **can be removed** with light guide
 - ◆ **Self shielding power**
- BG study
 - ◆ Understanding of **the source of BG**
 - ◆ Measuring **photon yield and its attenuation length**

100kg prototype

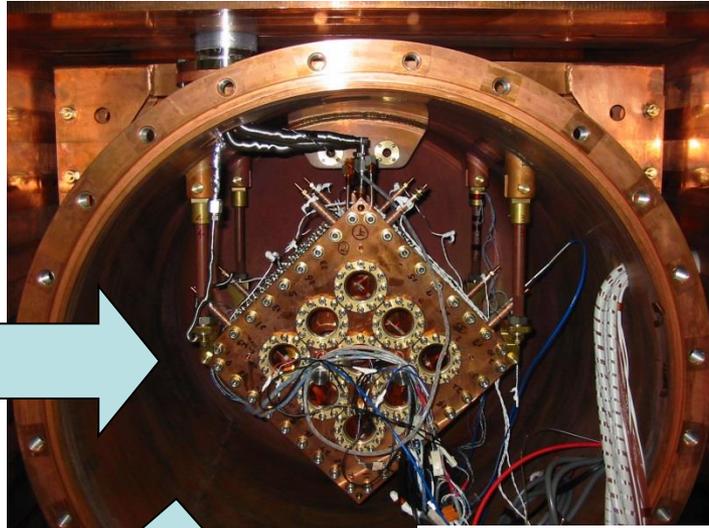
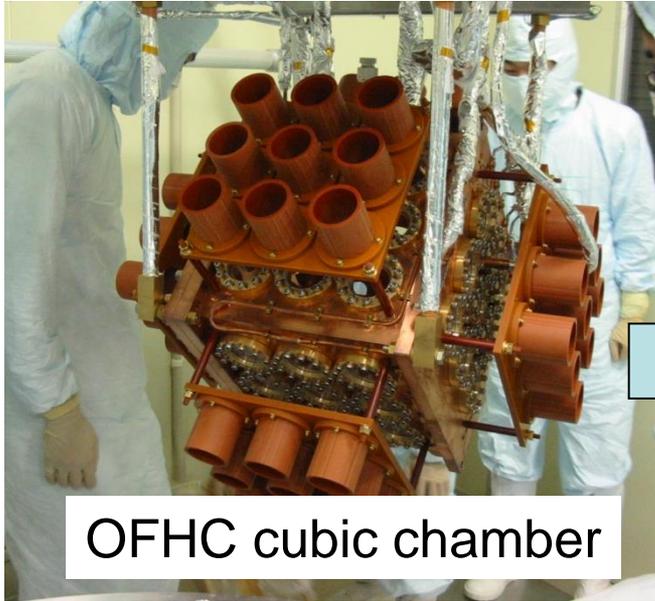


~30 cm cube
3 kg fiducial

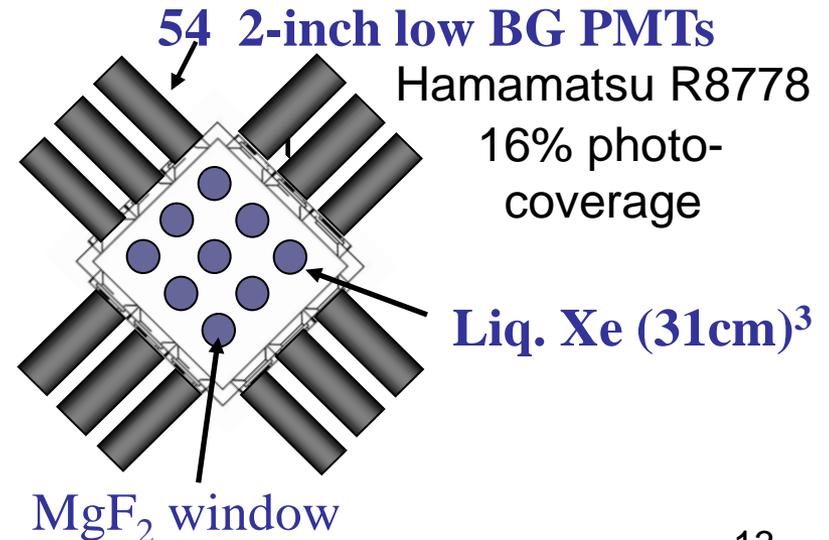
With **light guide**
version



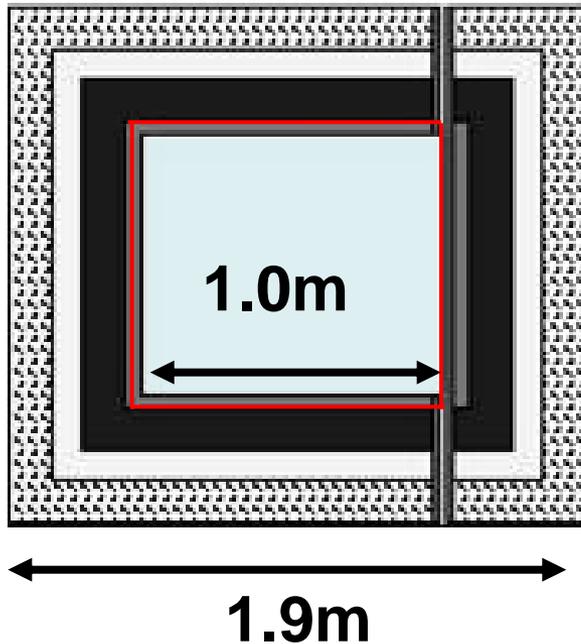
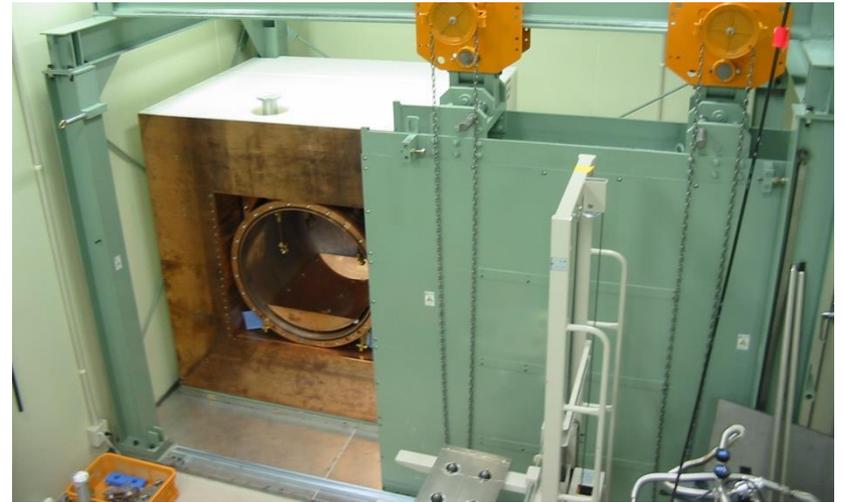
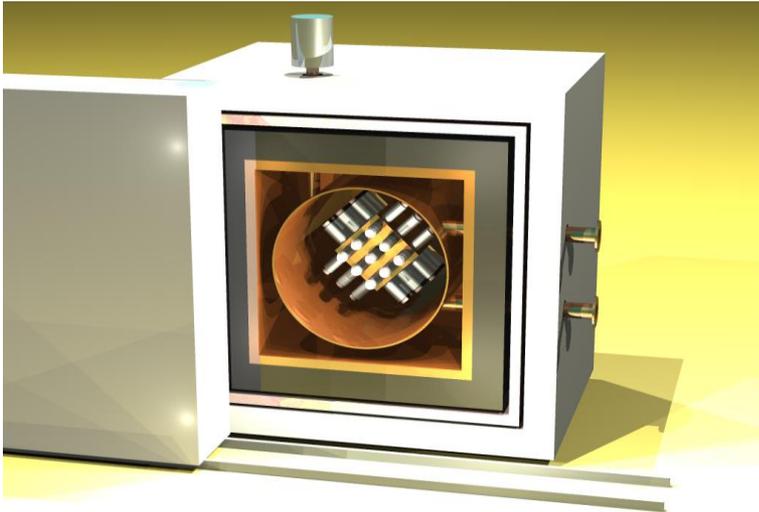
➤ 100 kg prototype detector



In the
Kamioka Mine
(near the Super-K)
2,700 m.w.e.



4 π shield with door



material	thickness
 Polyethylene	15cm
 Boron	5cm
 Lead	15cm
 EVOH sheets	30 μ m
 OF Copper	5cm

 Rn free air ($\sim 3\text{mBq/m}^3$)

➤ Progress so far

- 1st run (Dec. 2003)
 - ◆ Confirmed **performances of vertex & energy reconstruction**
 - ◆ Confirmed **self shielding power** for external γ rays
 - ◆ Measured **the internal background** concentration
- 2nd run (Aug. 2004)
 - ◆ Succeeded to **reduce Kr from Xe** by distillation
 - ◆ Photo electron yield is increased
 - ◆ Measured Rn concentration inside the shield
- 3rd run (Mar. 2005) with **light guide**
 - ◆ Confirmed **the miss fitting (only for the prototype detector) was removed**
 - ◆ Now, BG data is under analysis

➤ Vertex and energy reconstruction

Reconstruction is performed by PMT charge pattern (not timing)

Calculate PMT acceptances from various vertices by Monte Carlo.

Vtx.: compare acceptance map $F(x,y,z,i)$

Ene.: calc. from obs. p.e. & total accept.

$$\text{Log}(L) = \sum_{\text{PMT}} \text{Log}\left(\frac{\exp(-\mu)\mu^n}{n!}\right)$$

L: likelihood

$$\mu: \frac{F(x,y,z,i)}{\sum F(x,y,z,i)} \times \text{total p.e.}$$

n: observed number of p.e.

$F(x,y,z,i)$: acceptance for i-th PMT (MC)

VUV photon characteristics:

$$L_{\text{emit}}=42\text{ph/keV}$$

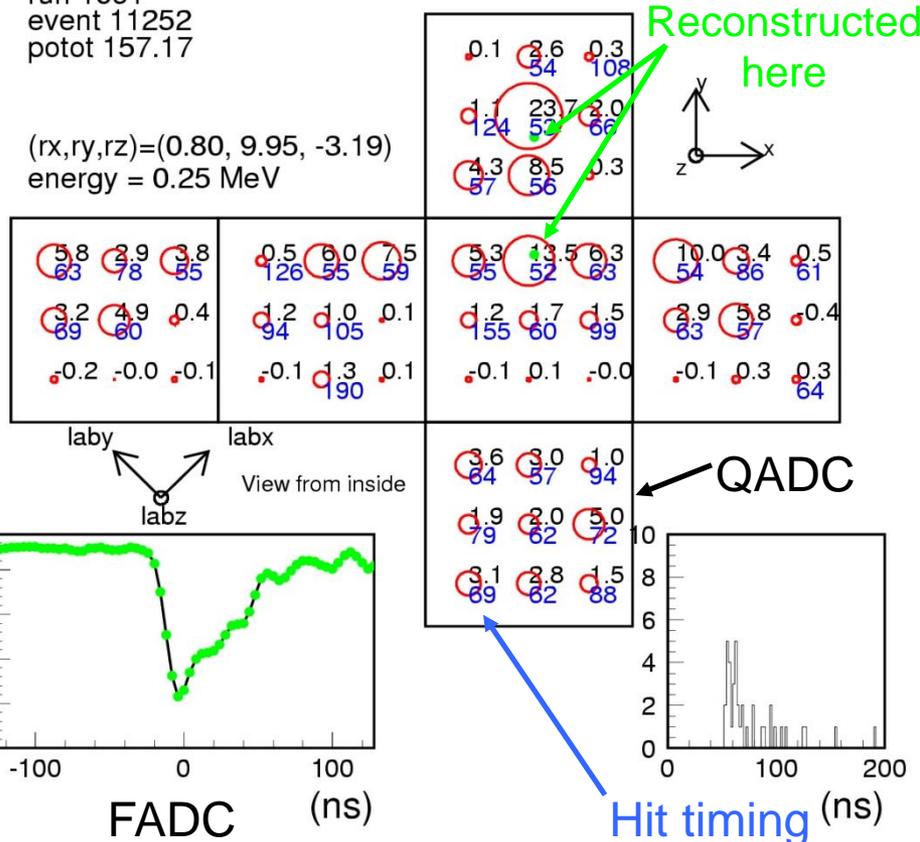
$$\tau_{\text{abs}}=100\text{cm}$$

$$\tau_{\text{scat}}=30\text{cm}$$

XMASS prototype detector

run 1091
event 11252
potot 157.17

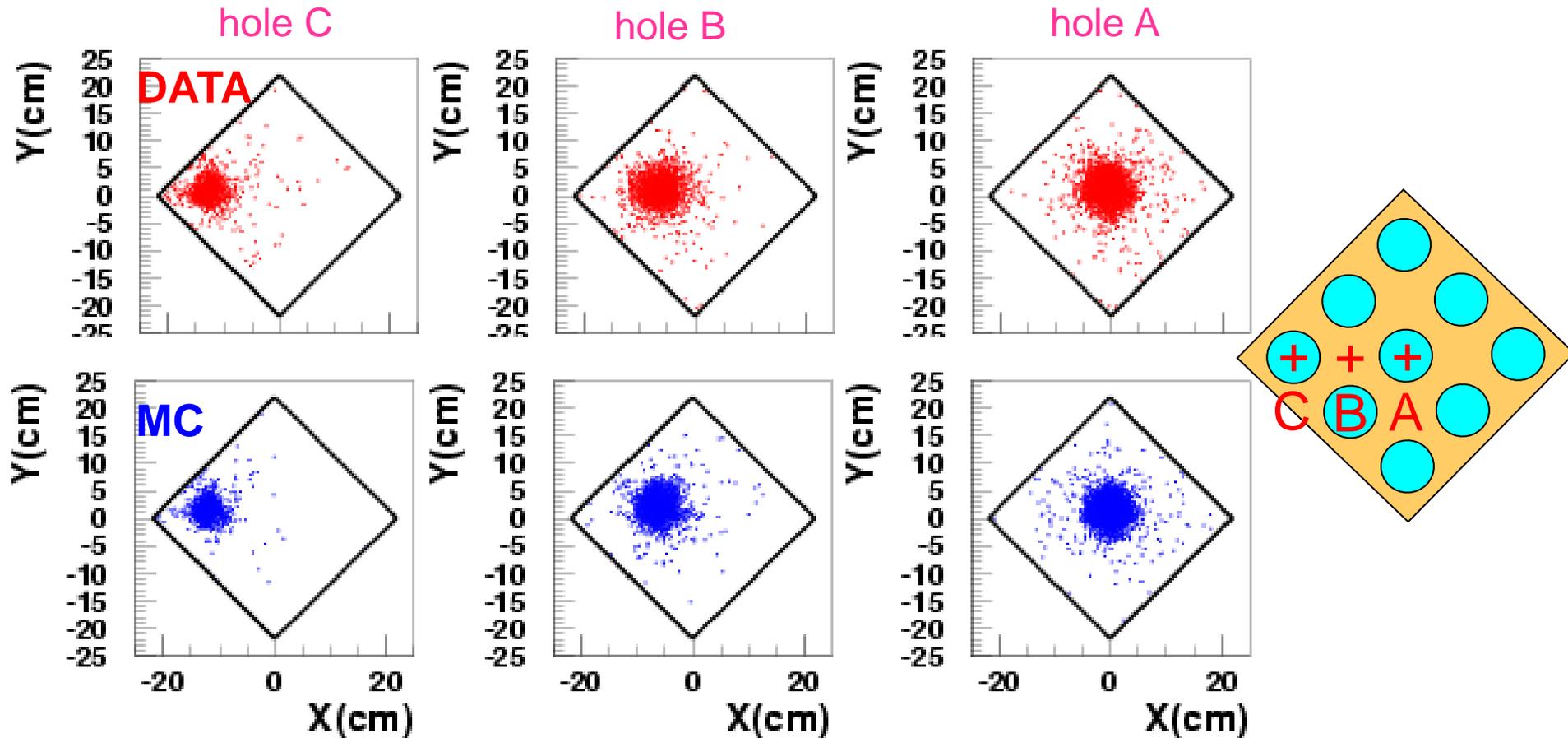
(rx,ry,rz)=(0.80, 9.95, -3.19)
energy = 0.25 MeV



=== Background event sample ===
QADC, FADC, and hit timing
information are available for analysis

1. Performance of the vertex reconstruction

Collimated γ ray source run from 3 holes (^{137}Cs , 662keV)

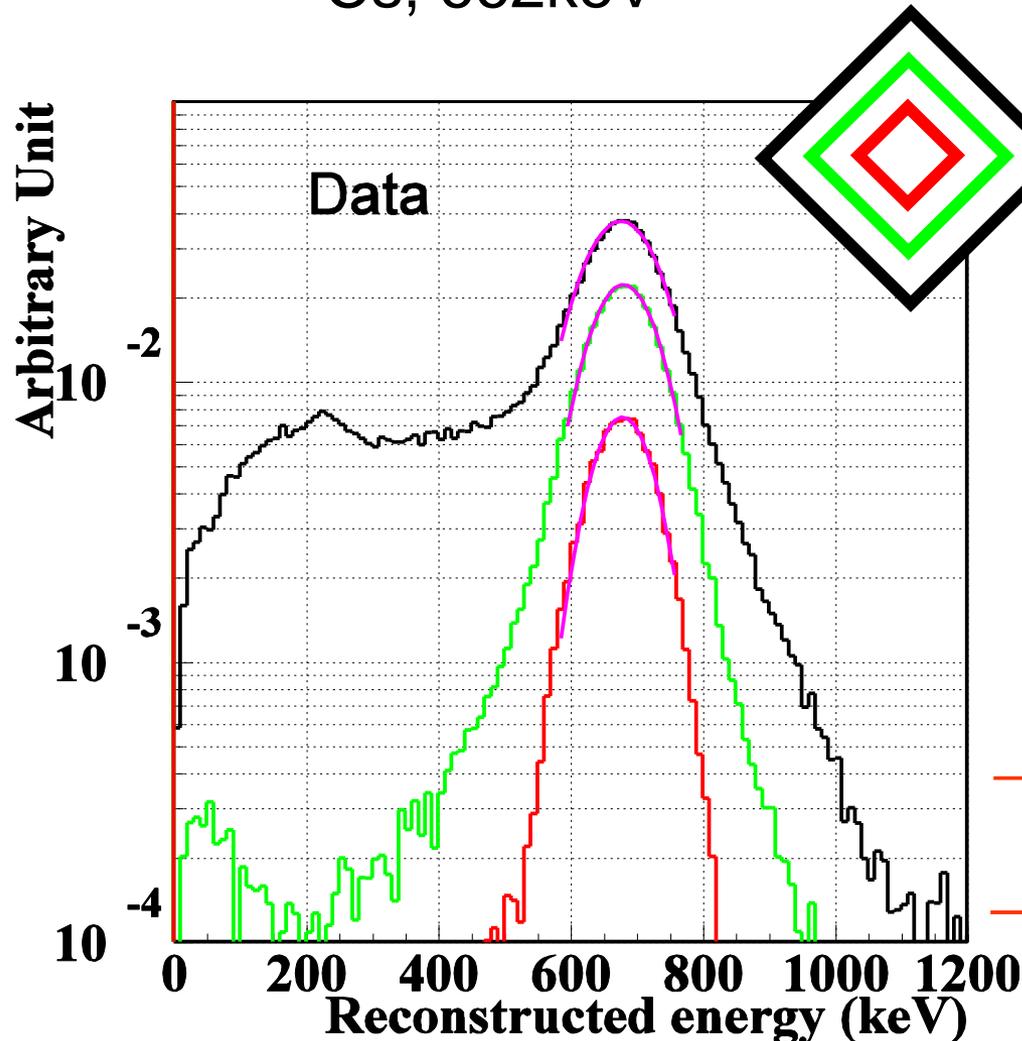


→ Vertex reconstruction works well

2. Performance of the energy reconstruction

Collimated γ ray source run from center hole

^{137}Cs , 662keV



- All volume
- 20cm FV
- 10cm FV

$\sigma=65\text{keV}@peak$
($\sigma/E \sim 10\%$)

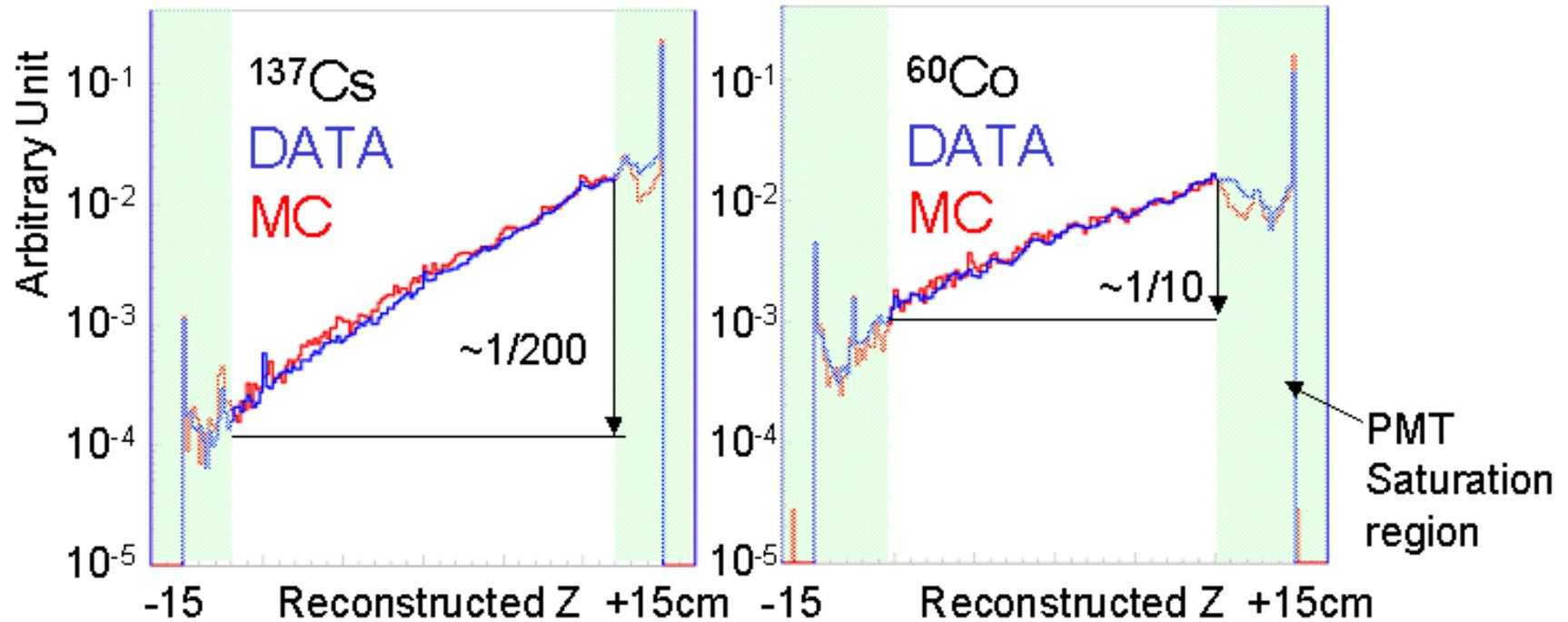
Similar peak position in
each fiducial.

→ No position bias

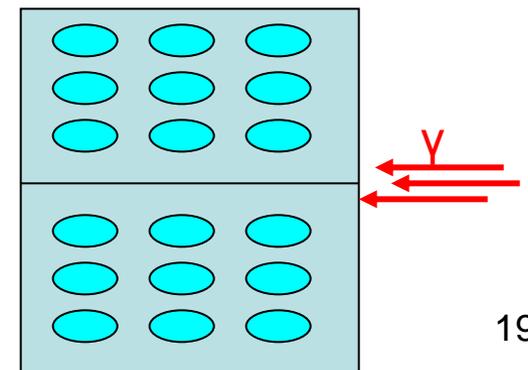
→ Energy reconstruction
works well

➤ Demonstration of self shielding effect

z position distribution of the collimated γ ray source run



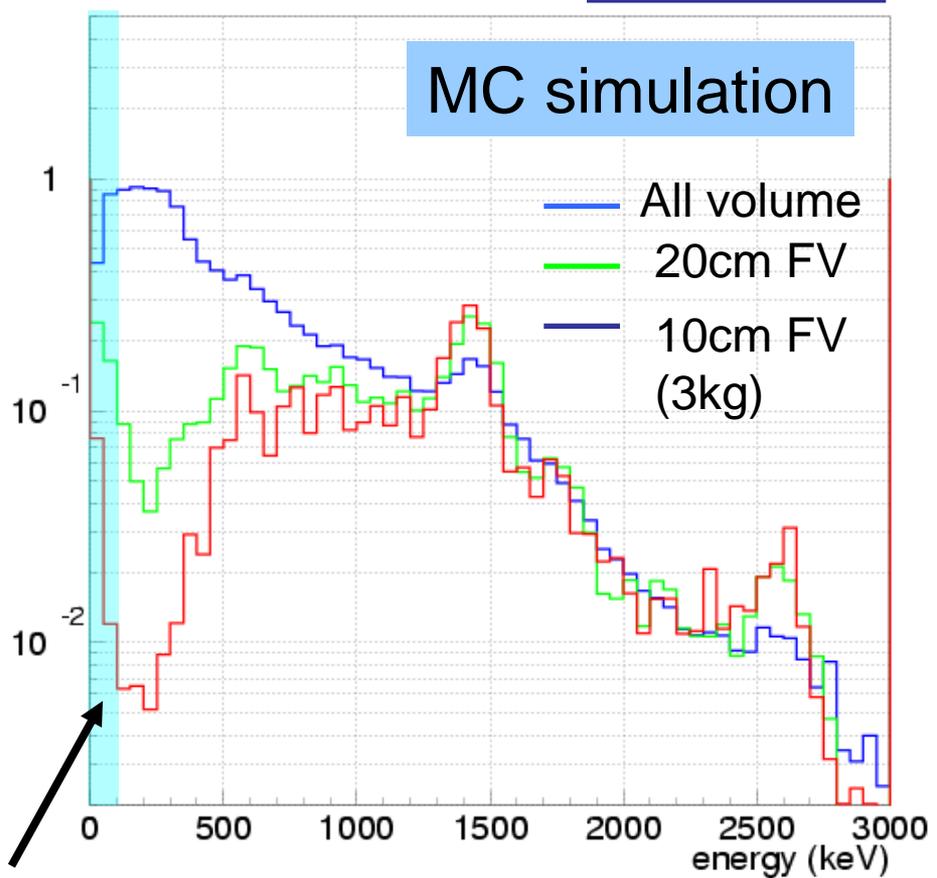
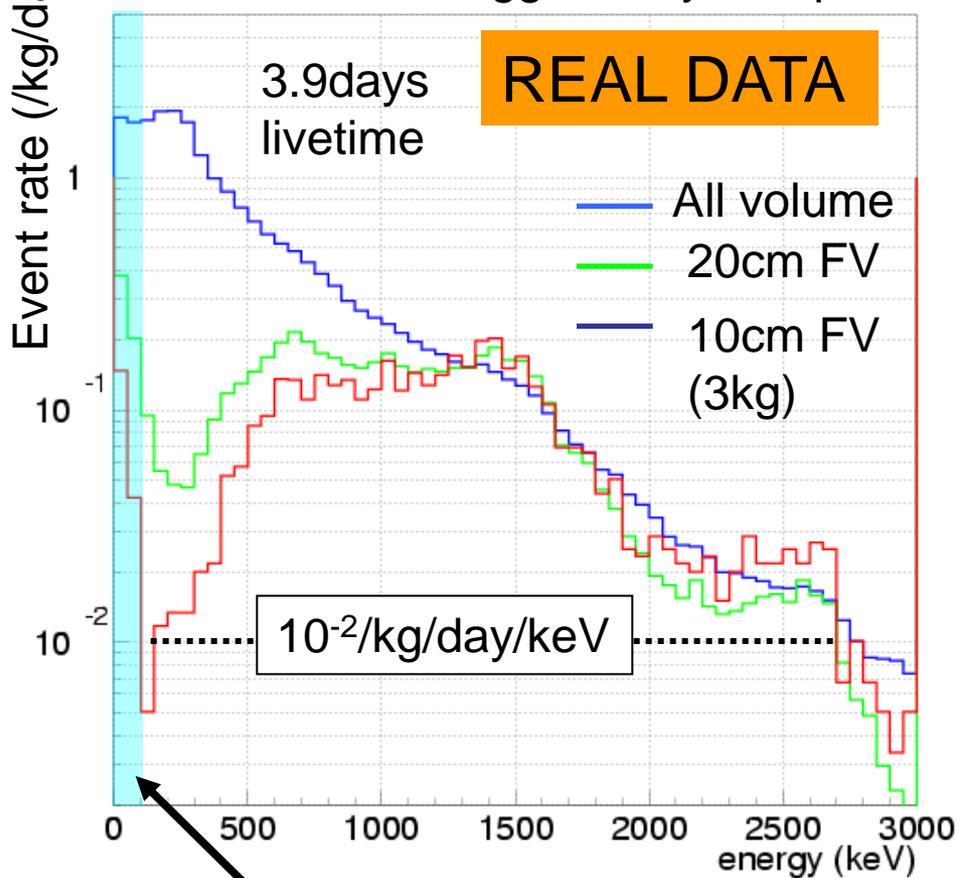
→ Data and MC agree well



Shelf shielding for real data and MC

Aug. 04 run
preliminary

~1.6Hz, 4 fold, triggered by ~0.4p.e.



Miss-reconstruction due to dead-angle region from PMTs.

- **Good agreement** (< factor 2)
- **Self shielding effect** can be seen clearly.
- **Very low background** (10^{-2} /kg/day/keV @ 100-300 keV)

➤ Internal backgrounds in liq. Xe were measured

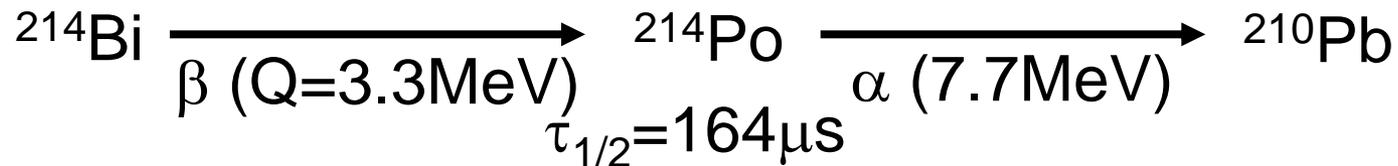
Main sources in liq. Xe are **Kr**, **U-chain** and **Th-chain**

- **Kr** = 3.3 ± 1.1 ppt (by mass spectrometer)

→ Achieved by distillation

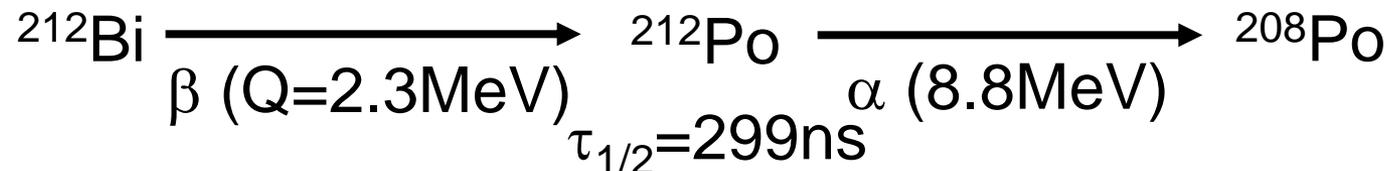
- **U-chain** = $(33 \pm 7) \times 10^{-14}$ g/g (by prototype detector)

Delayed coincidence search (radiation equilibrium assumed)



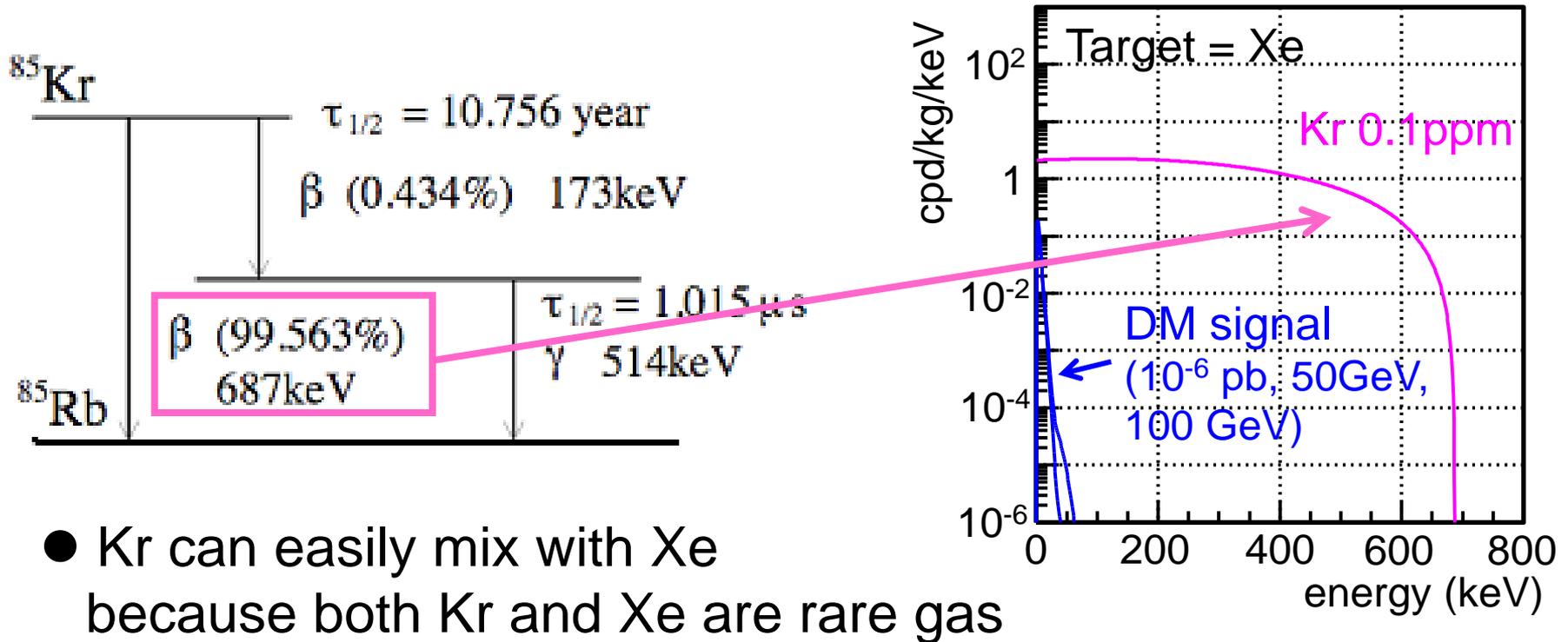
- **Th-chain** < 23×10^{-14} g/g (90%CL) (by prototype detector)

Delayed coincidence search (radiation equilibrium assumed)



Kr concentration in Xe

- ^{85}Kr makes BG in low energy region



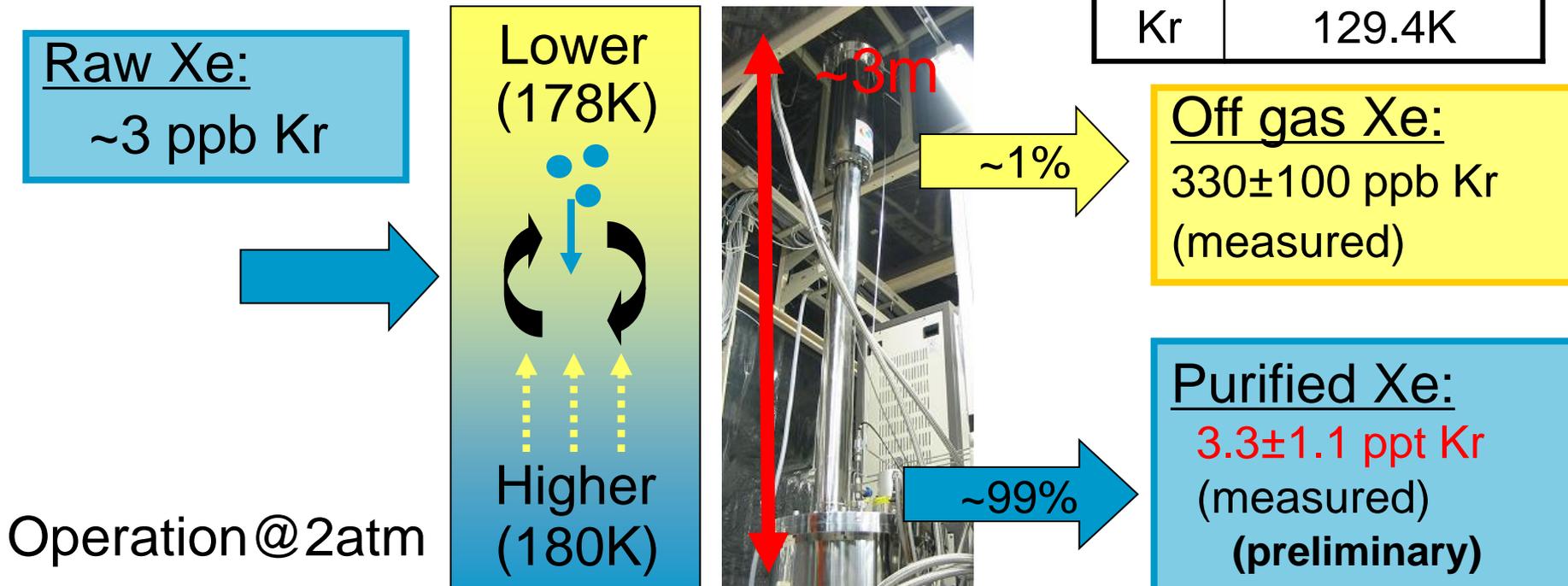
- Kr can easily mix with Xe because both Kr and Xe are rare gas
- Commercial Xe contains a few ppb Kr

Xe purification system

- XMASS succeeds to reduce Kr concentration in Xe from ~3[ppb] to $3.3(\pm 1.1)$ [ppt] with one cycle (~1/1000)

- Processing speed : 0.6 kg / hour
- Design factor : **1/1000 Kr** / 1 pass
- Purified Xe : Off gas = 99:1

	Boiling point (@2 atm)
Xe	178.1K
Kr	129.4K



Summary of BG measurement

Now (prototype detector)

Goal (800kg detector)

- γ ray BG $\sim 10^{-2}$ cpd/kg/keV

$\xrightarrow{1/100}$

10^{-4} cpd/kg/keV

→ Increase volume for self shielding

→ Decrease radioactive impurities in PMTs ($\sim 1/10$)

- $^{238}\text{U} = (33 \pm 7) \times 10^{-14}$ g/g

$\xrightarrow{1/33}$

1×10^{-14} g/g

→ Remove by filter

- $^{232}\text{Th} < 23 \times 10^{-14}$ g/g (90% C.L.)

$\xrightarrow{1/12}$

2×10^{-14} g/g

→ Remove by filter (Only upper limit)

- Kr = 3.3 ± 1.1 ppt

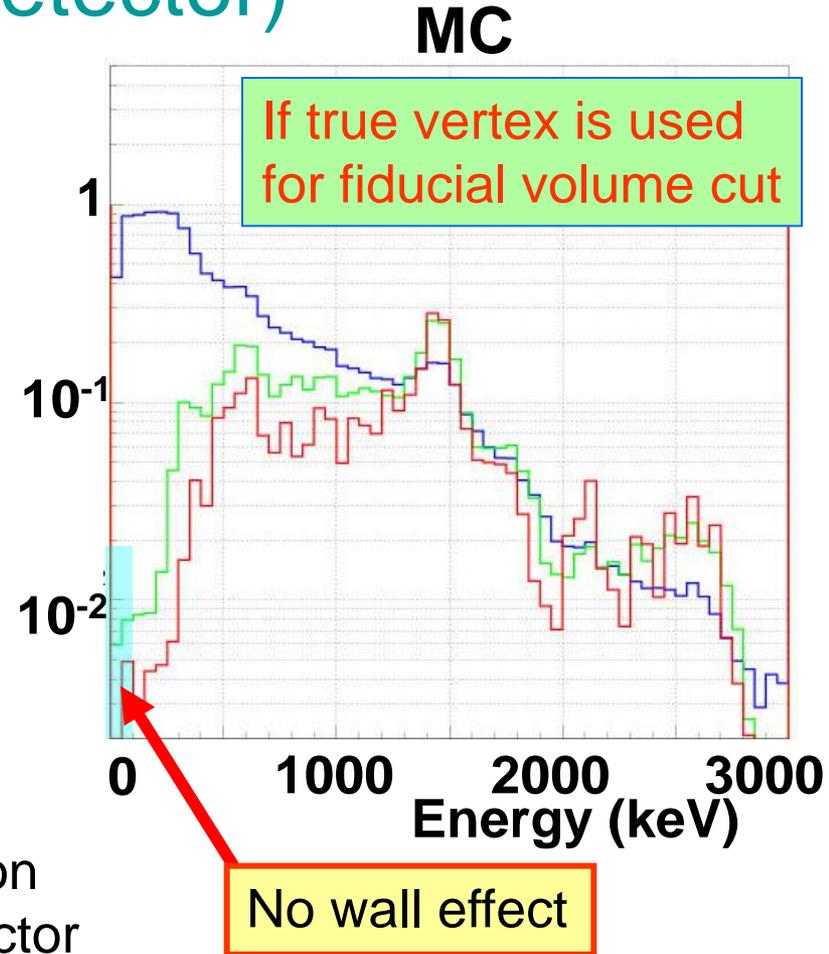
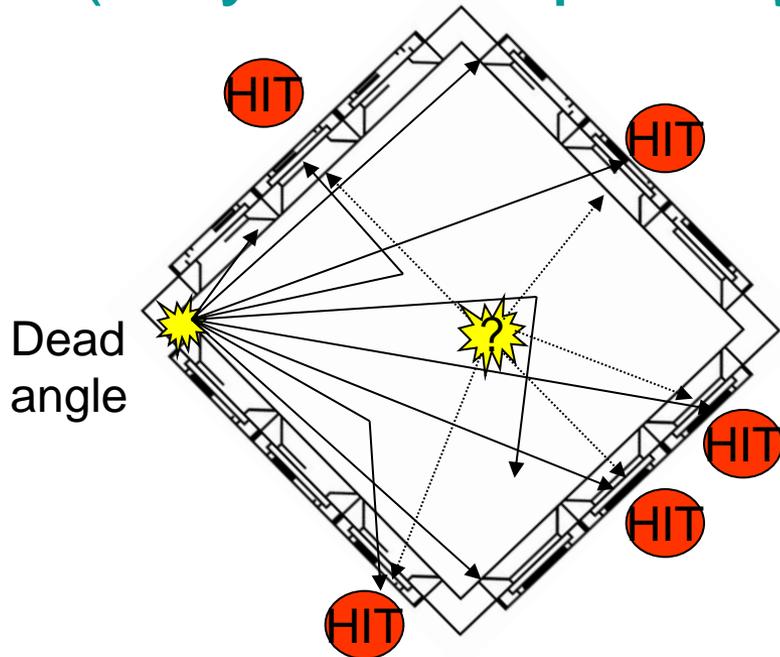
$\xrightarrow{1/3}$

1 ppt

→ Achieve by 2 purification pass

Very near to the target level!

➤ Remaining problem: wall effect (only for the prototype detector)

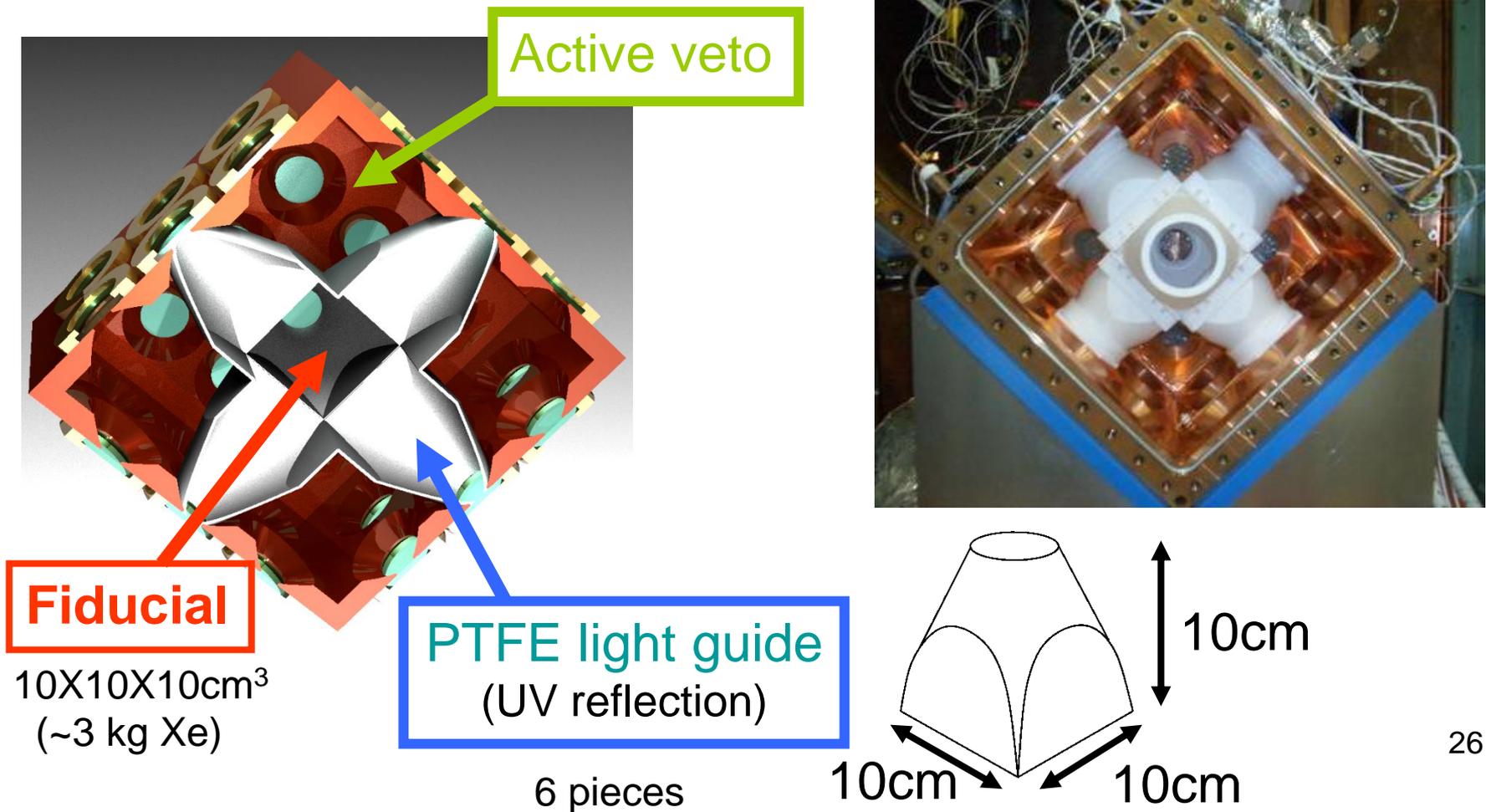


- Scintillation lights at the dead angle from PMTs give quite uniform 1 p.e. signal for PMTs, and this cause miss reconstruction as if the vertex is around the center of detector

➔ This effect does not occur with the sphere shape 800 kg detector

➤ Prototype detector with light guide

Purpose: **remove the wall effect and understand the source of BG in the DM region**



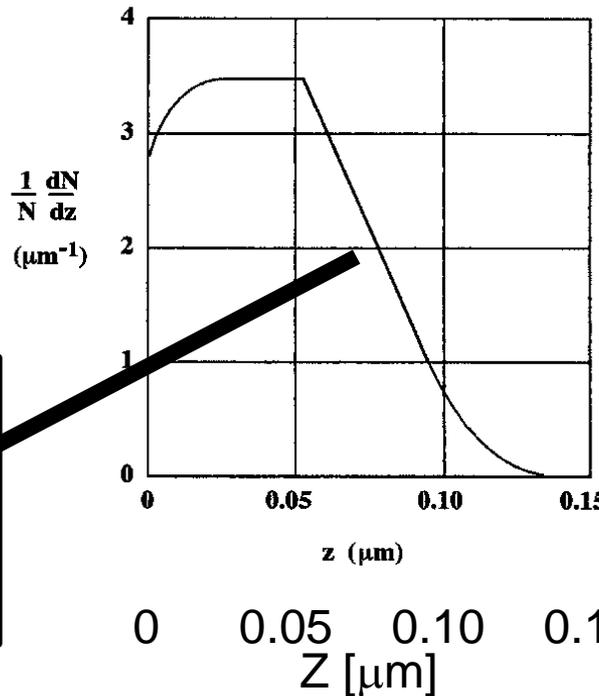
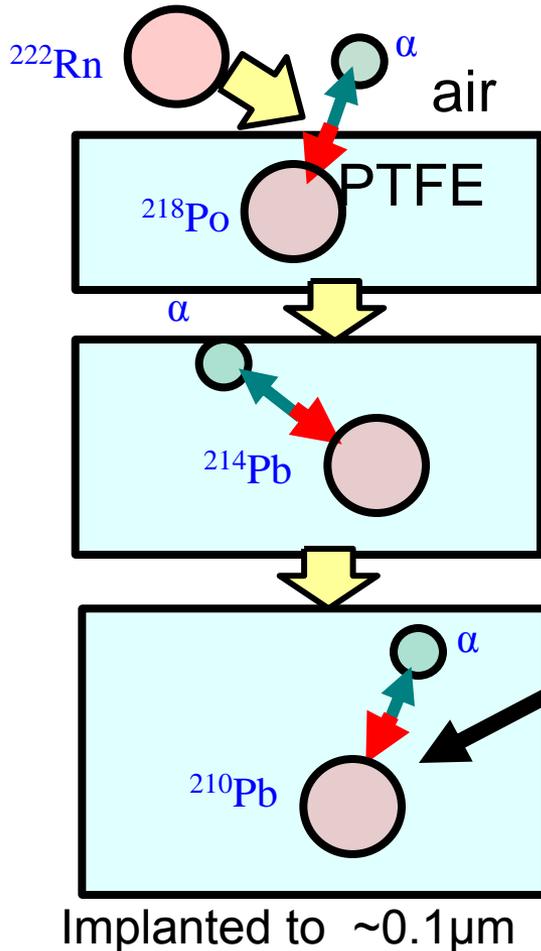
Light guide setup

- Edging of PTFE surfaces

^{222}Rn decays (^{210}Pb β , 64 keV endpoint) implanted in PTFE surfaces might make the dominant BG



We edged the PTFE inside $\sim 10\mu\text{m}$



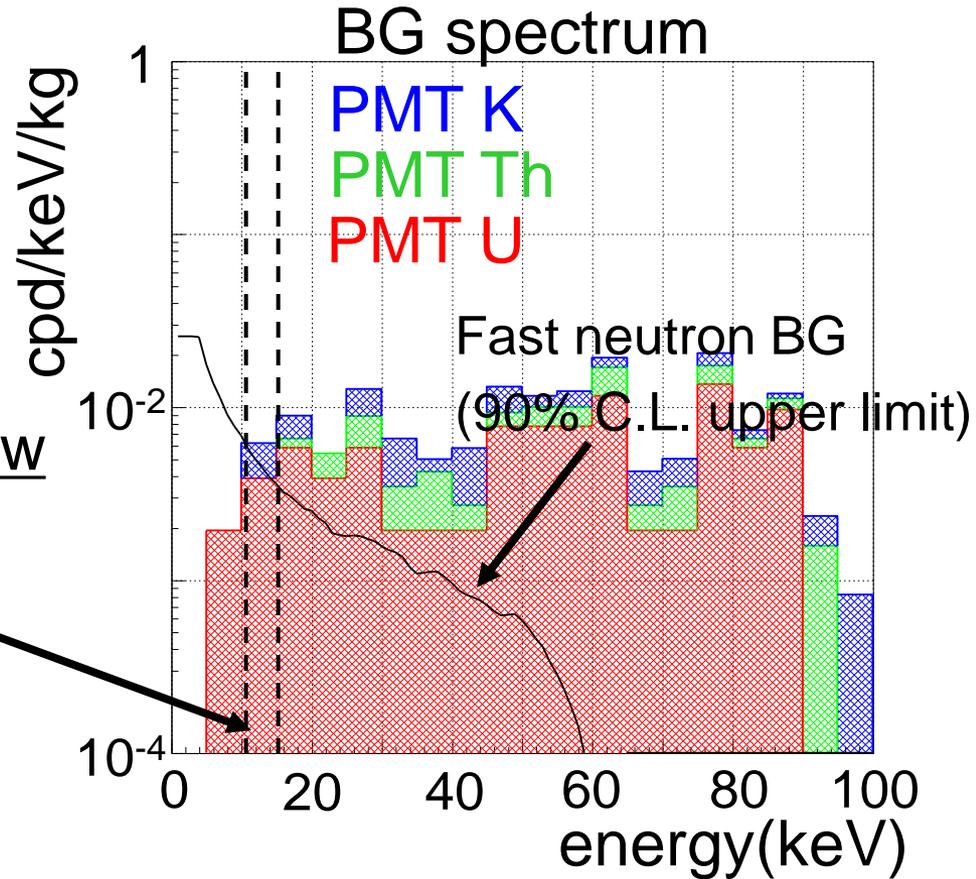
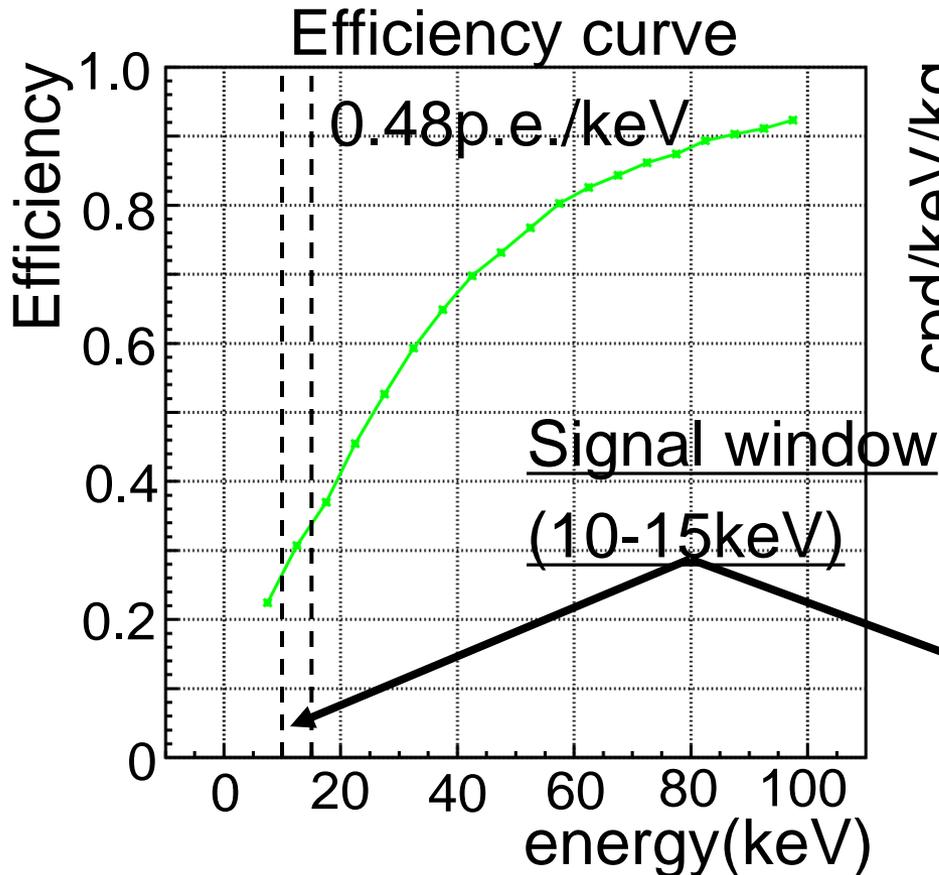
◀ N.J.T. Smith et al., Phys. Lett. B 485 (2000) 9

Position distribution of ^{210}Pb (in NaI)

Recoil process implants
30% of the original
 surface Rn decays

Expected BG spectrum

- MC simulation was done with GEANT3



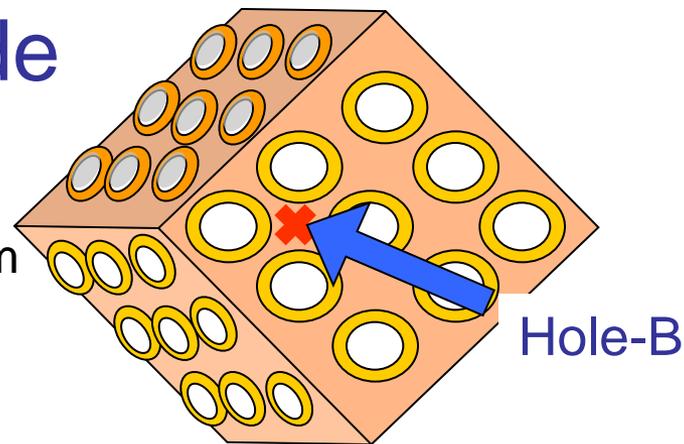
Efficiency ~30% @ 10keV

Expected BG ~ 10^{-2} cpd/keV/kg

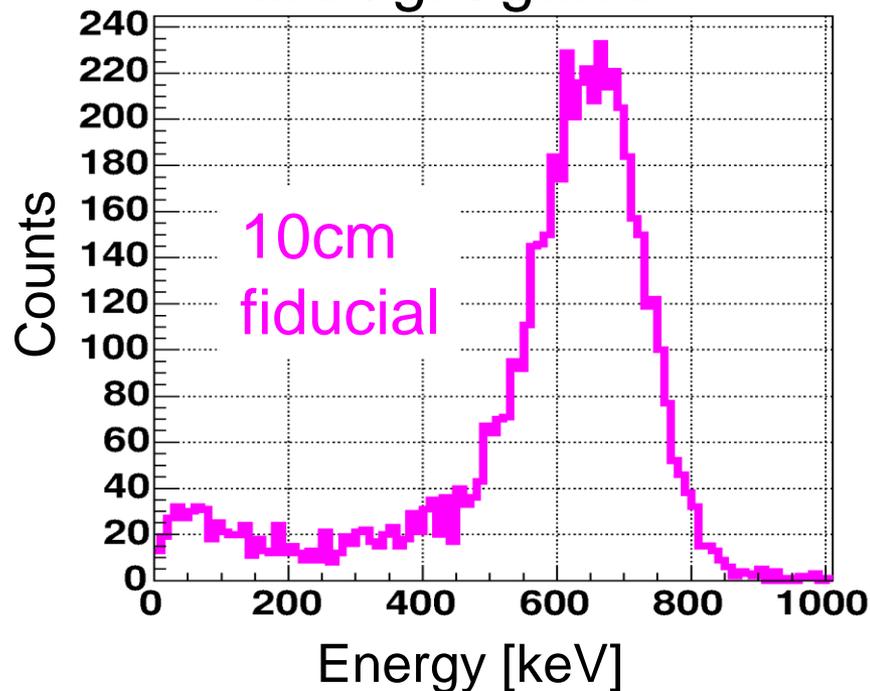
→ Very low BG ~ 10^{-2} cpd/keV/kg @ <100keV

Result 1: comparing the data taken with and without light guide

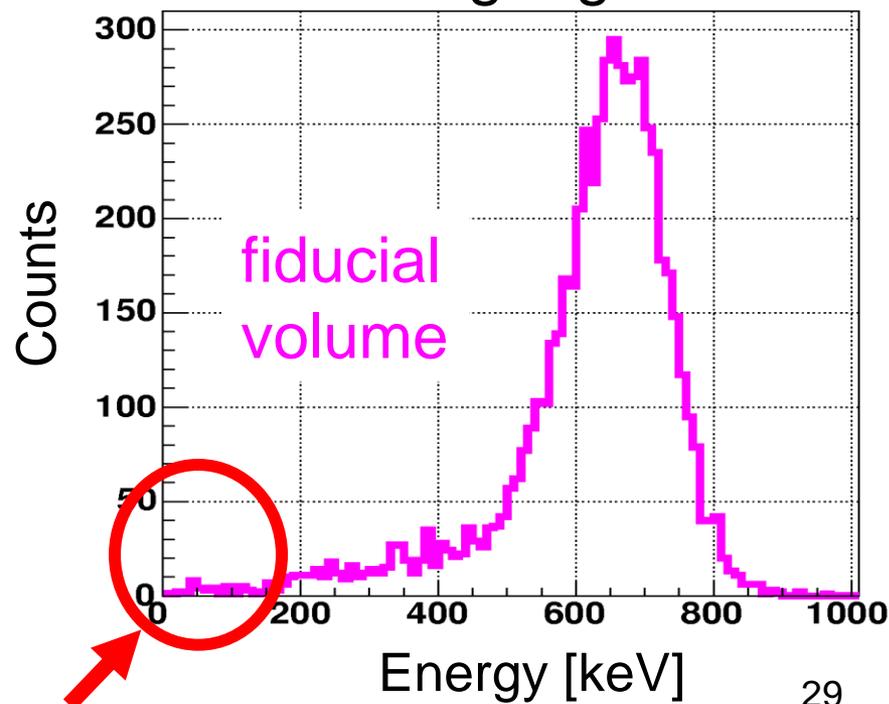
Collimated γ ray source run from hole-B (137Cs, 662keV)



● w/o light guide

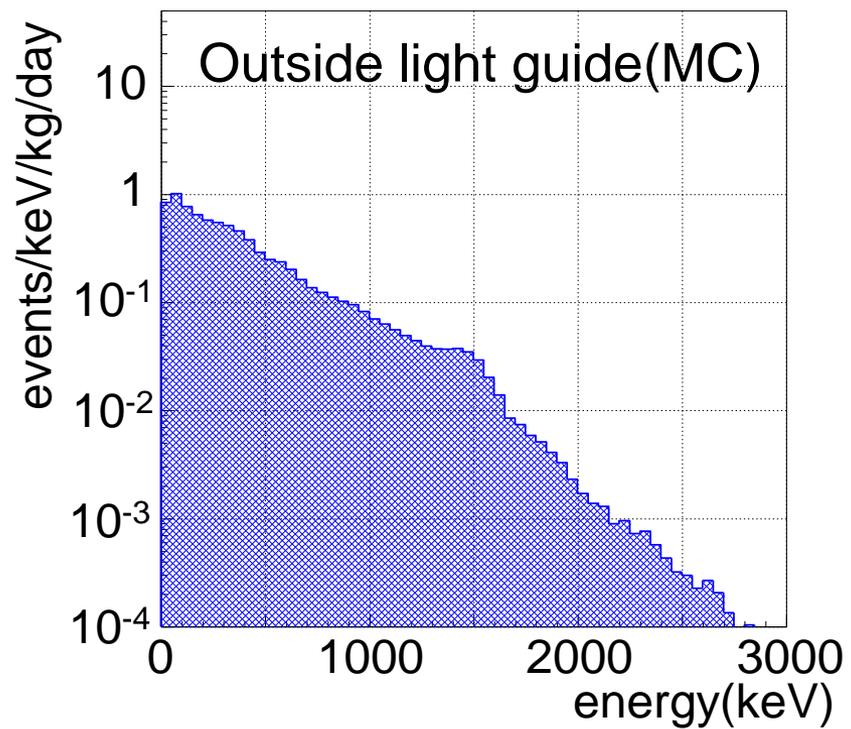
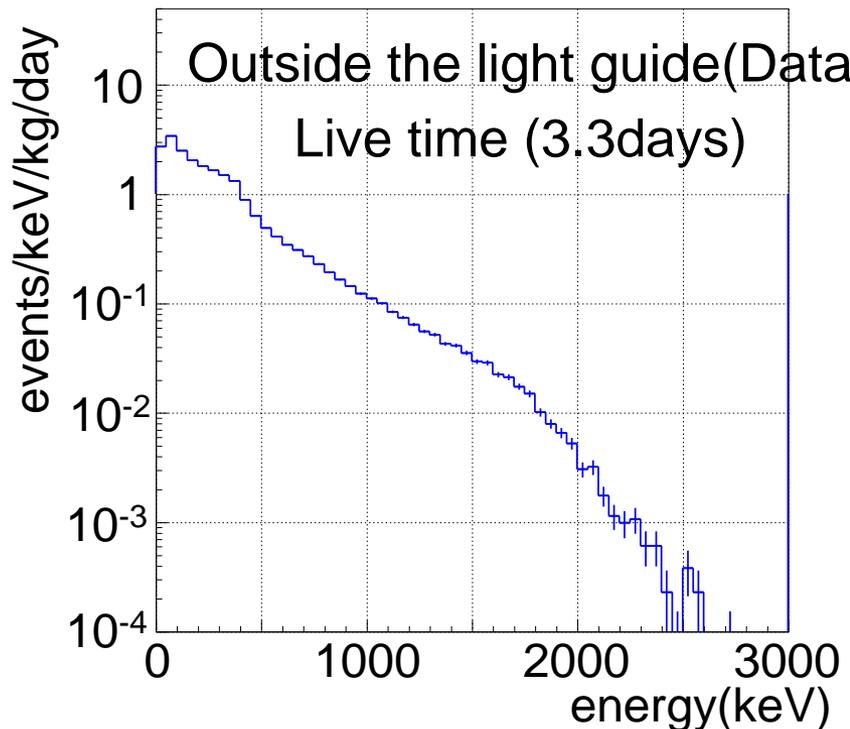


● with light guide



Reduce the events due to the wall effect

Result 2: Obtained energy spectrum outside the light guide



- **Good agreement** (< factor 2)
- Trigger rate is same as the measurement without guide (Aug. 2004)

3. Summary

- XMASS experiment:
Multi purpose low-background experiment with large mass liq. Xe
- 800 kg detector:
Designed for dark matter shearch mainly, and **10^2 improvement of sensitivity** above existing experiments is expected
- R&D with the 100 kg prototype detector
Most of the **performances required for 800 kg detector are confirmed**